

# Using calibration from 3D simulations in stellar evolution calculations

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# Significance of convection treatment

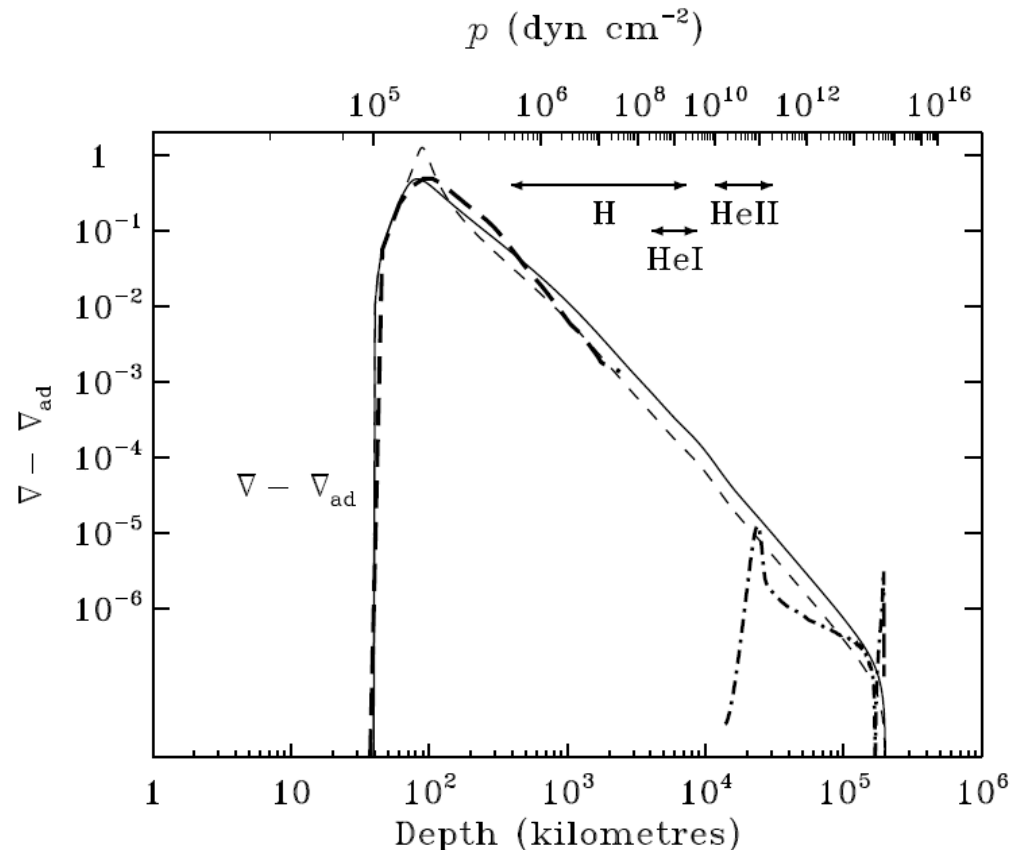
- Defines behaviour of superadiabatic gradient in top of convection zone
- **Given atmospheric properties** defines specific entropy (adiabat) in nearly adiabatic part of convection zone
- Hence needs consistency between atmosphere and mixing-length calibration

# Significance of convection treatment

- Defines behaviour of superadiabatic gradient in top of convection zone

- **Given atmospheric entropy (adiabat) convection zone**

— Böhm-Vitense  
- - - Canuto & Mazzitelli  
- - - Averaged simulation

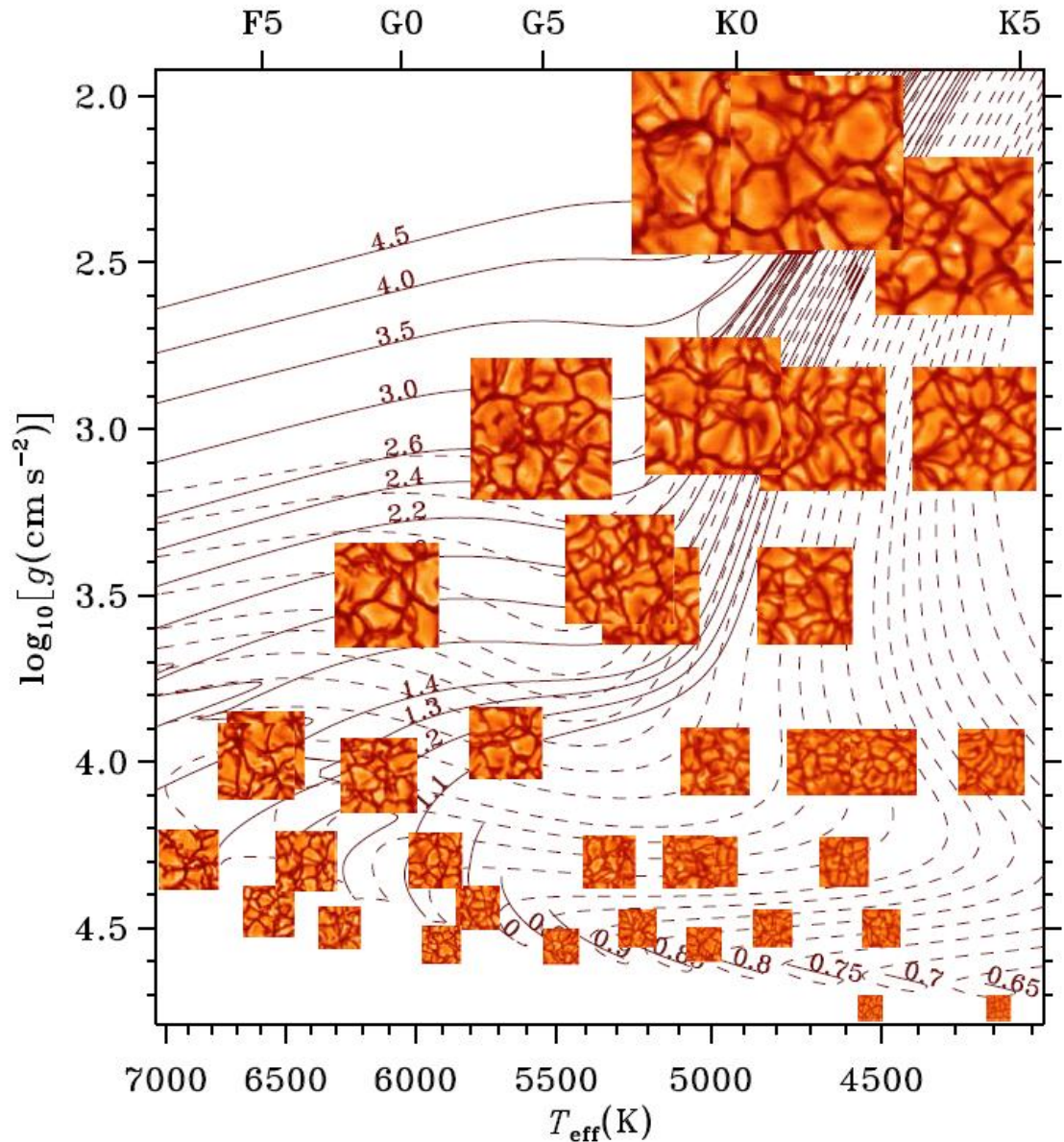


# Calibration from 3D simulation

- Define atmospheric properties in terms of  $T(\tau)$
- Determine  $\alpha$  such that structure in 1D envelope model matches deep (nearly adiabatic) part of the average simulation
- Interpolate to model  $T_{\text{eff}}$  and  $\log g$  in grid of simulations

# Simulations

Solar composition:  
 $X = 0.737, Z = 0.018$

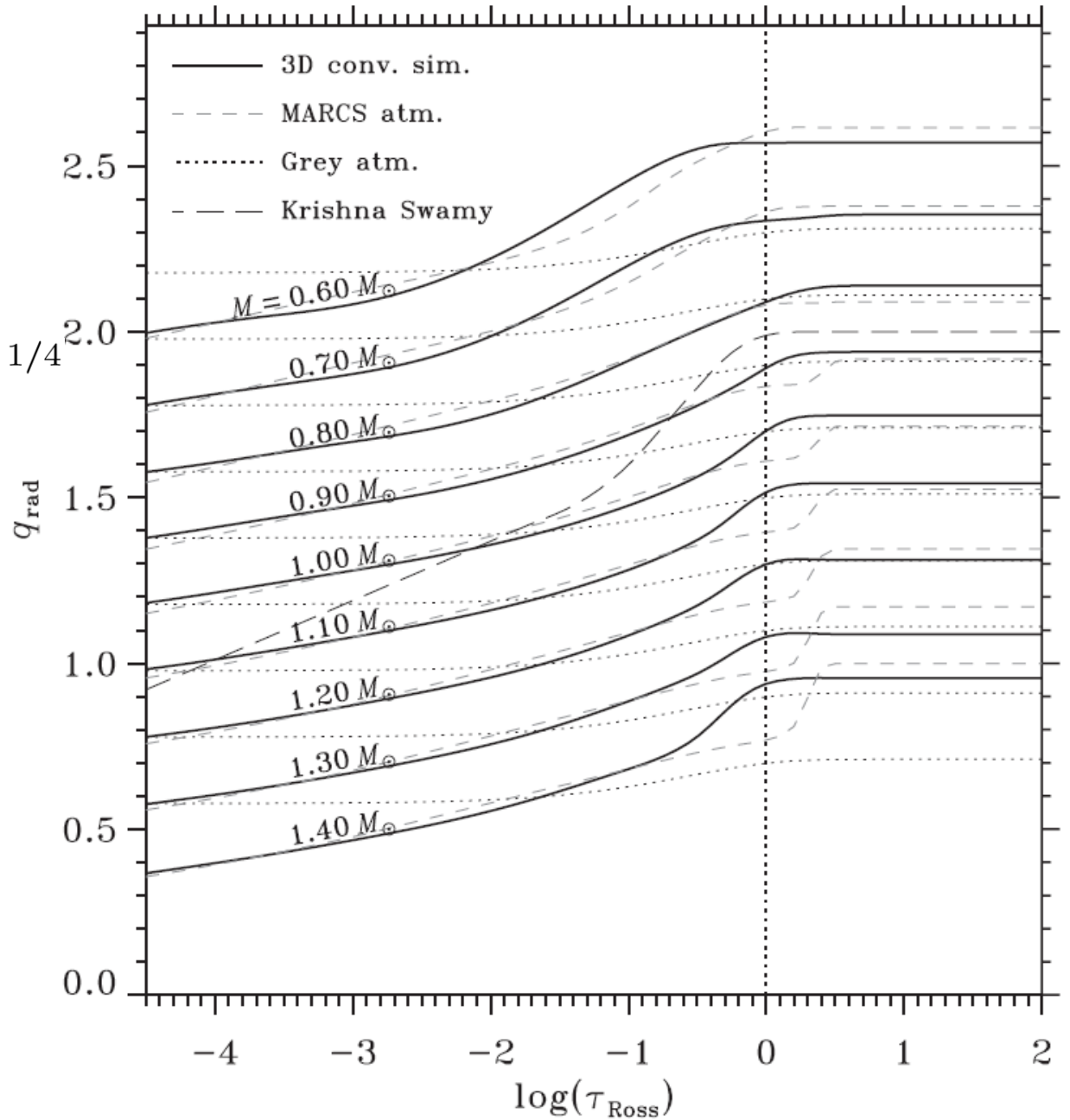


Trampedach et al. (2014;  
MNRAS 445, 4366)

# $T(\tau)$

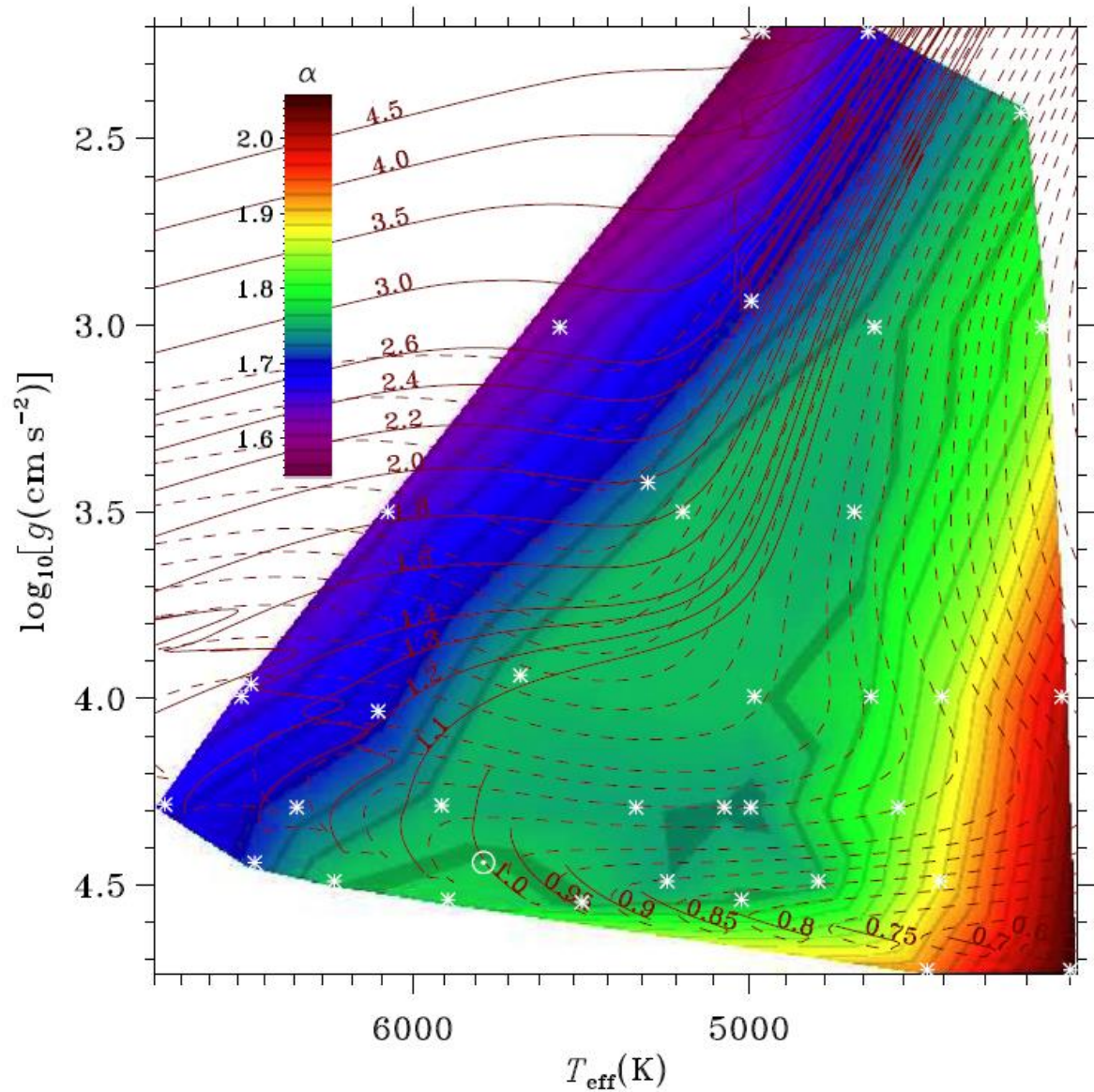
Characterize by

$$T(\tau) = T_{\text{eff}} \left[ \frac{3}{4} (q(\tau) + \tau) \right]^{1/4}$$



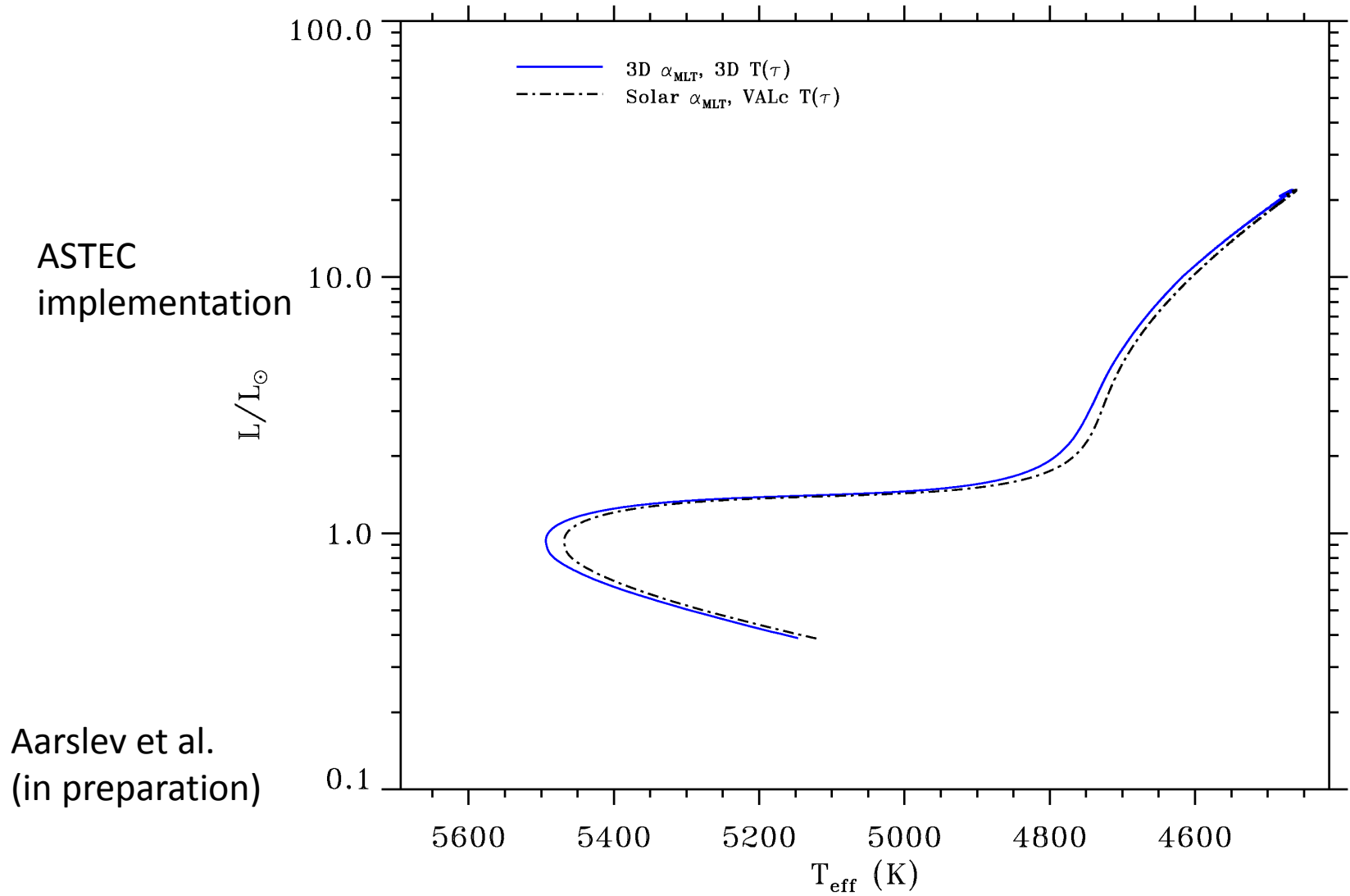
Trampedach et al. (2014;  
MNRAS 442, 805)

# Calibrated $\alpha$



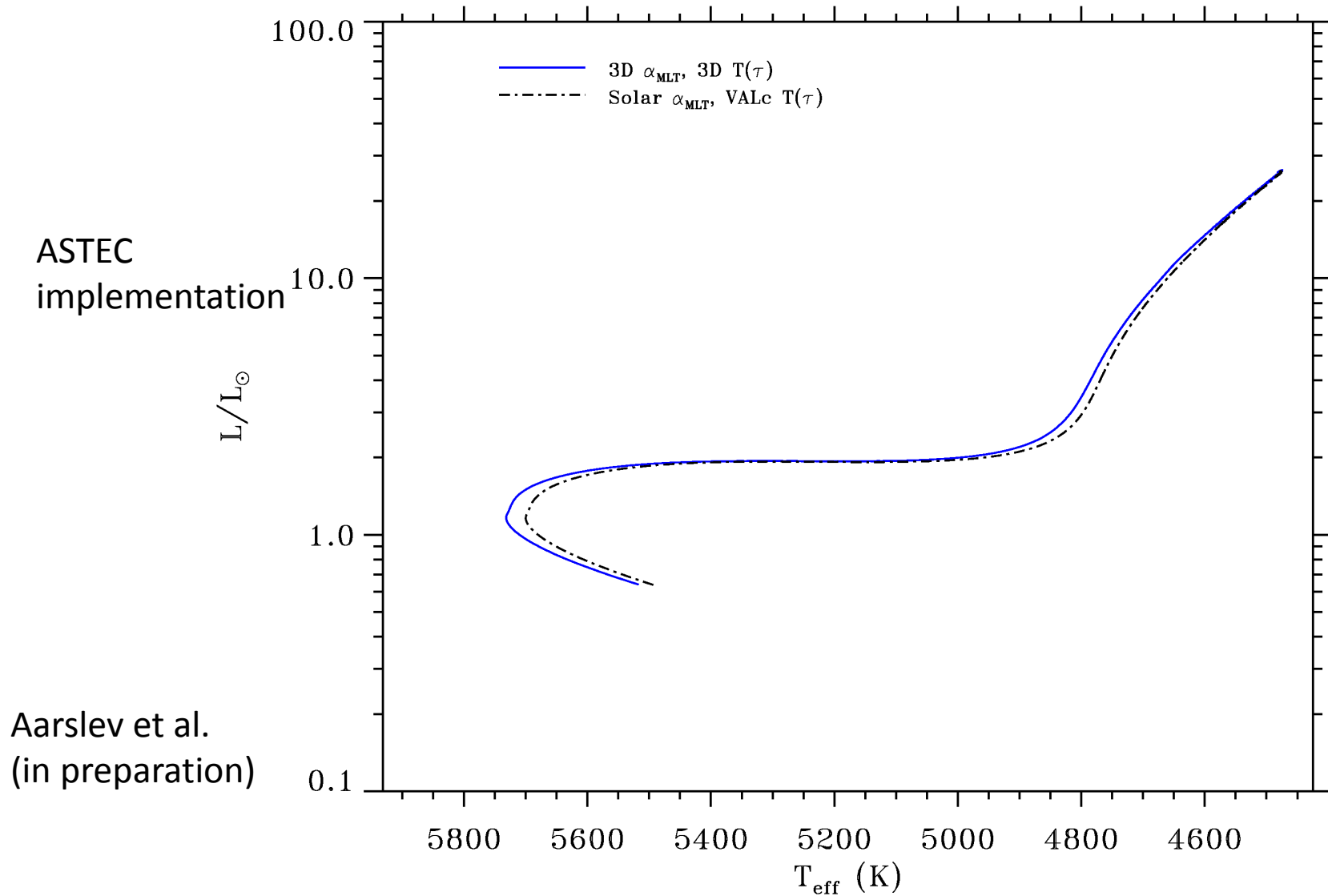
Trampedach et al. (2014;  
MNRAS 445, 4366)

# 0.9 M<sub>⊙</sub>

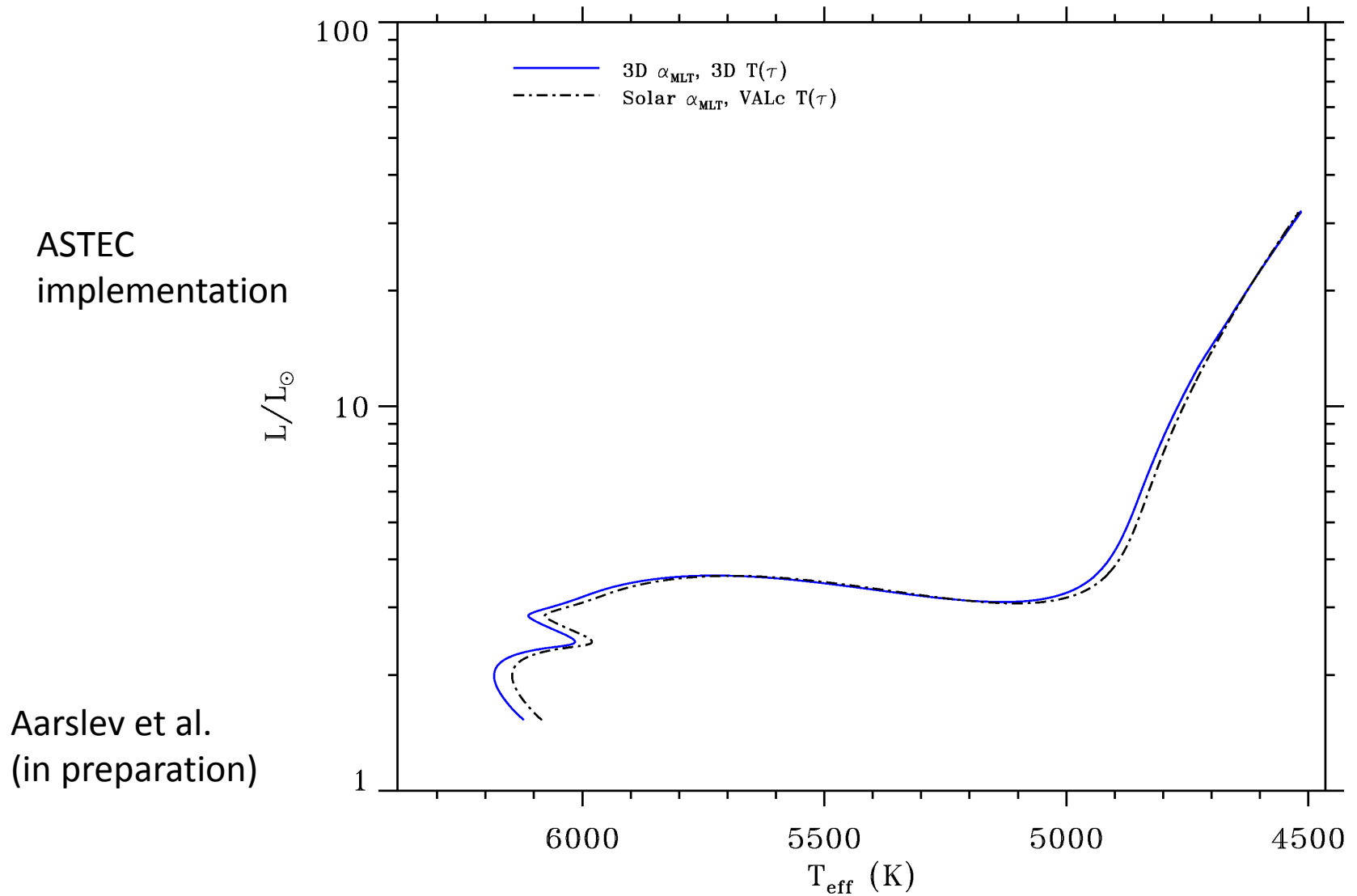




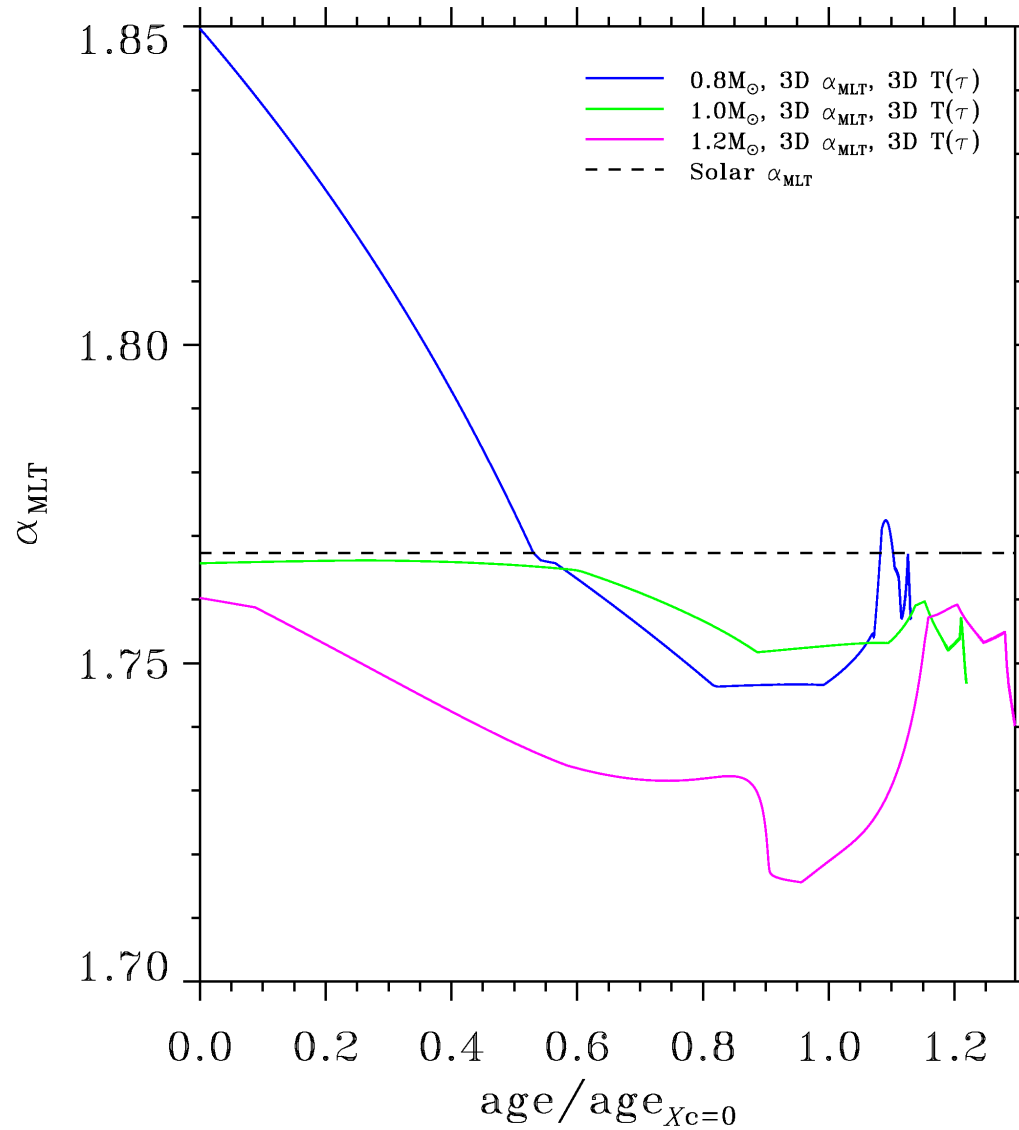
# 1.0 M<sub>⊙</sub>



# 1.2 M<sub>⊙</sub>



# Variation of $\alpha$



# Things to come

- Independent application with GARSTEC models (Jakob Mosumgaard et al.)
- Use of existing Stagger grid (Magic et al.)
- Extension of Trampedach grid to a range of compositions
- Implementation of calibrated turbulent pressure
- Direct fit to interpolated simulations